

Supersymmetric Electromagnetic Waves on Giant Gravitons

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Introduction

- Supersymmetric states have played a crucial role in the development of string theories ...
... by helping to uncover and substantiate important aspects about dualities.
- This is especially true of the strong-weak dualities.
- In the context of AdS/CFT the program of matching the BPS spectra on both sides of the duality is still not completed.
- ... will be useful in the verification of AdS/CFT in its BPS sectors.
- ... may help in accounting for the entropy of the extremal black holes in $AdS_5 \times S^5$.

- In the context of AdS/CFT a significant class of BPS states consists of the so-called giant graviton states of type IIB string theory on $AdS_5 \times S^5$.
- These are supposed to be dual to the R-charged BPS states of the $\mathcal{N} = 4$ $SU(N)$ SYM on S^3 under the AdS/CFT dictionary.
- A lot of progress has been made in counting the R-charged states on the CFT side.
- For instance it is known that the degeneracy of the 1/8-BPS states of $\mathcal{N} = 4$ $U(N)$ SYM is identical to that of N particles in 3 bosonic and 2 fermionic oscillators. [Kinney, Maldacena, Minwalla, Raju]

- A part of this answer has been recovered from the bulk side.

[Biswas, Gaiotto, Lahiri, Minwala; Mandal-NVS]

- ... this is achieved by starting with classical BPS solutions of D3-branes in $AdS_5 \times S^5$ and quantizing their solution spaces.
- However, we do not yet know the *full* set of BPS solutions of D3-branes with in a fixed susy sector.
- Thus it is important to improve our knowledge of classical solutions of D3-branes in $AdS_5 \times S^5$.

- States in $AdS_5 \times S^5$ are labeled by the eigen values of the commuting set of isometries ($SO(2, 4) \times SO(6)$)

$$(E, S_1, S_2; J_1, J_2, J_3)$$

- BPS states preserve at least 2 susy out of the 32 of $AdS_5 \times S^5$, and satisfy

$$E = |S_1| + |S_2| + |J_1| + |J_2| + |J_3|.$$

- BPS D3-brane configurations in $AdS_5 \times S^5$ fall into two categories:

Giant

A D3-brane on a 3-surface in S^5 , point-like in AdS_5 that moves along a time-like geodesic.

Dual-Giant

A D3-brane on a 3-surface in AdS_5 , point-like in S^5 and moving along a maximal circle.

- Several examples are known:
characterized by their susy and non-zero charges.

- There are three sectors with (at least) 4 susy:

$$(J_1, J_2, J_3), (S_1, J_1, J_2), (S_1, S_2, J_1).$$

- Part of the (J_1, J_2, J_3) sector has two descriptions:
 - (i) Mikhailov giants [Mikhailov] (ii) spinning dual-giants [Mandal-NVS]
- Part of the (S_1, S_2, J_1) sector also has two descriptions
 - (i) Spinning giants [Mandal-NVS] (ii) Wobbling dual-giants [Ashok-NVS]
- In all these solutions only the transverse scalars of D3-branes excited.
- Mikhailov giants and Wobbling dual-giants are the full set of solutions with the scalars turned on in their susy sectors. [Ashok-NVS]

- In general one expects solutions with transverse scalars, the gauge fields and fermions turned on.
- Some isolated examples of BPS gauge field configurations on round S^3 giants (1/2-BPS) were known [Kim, Lee; Sinha, Sonner]
- In this talk we address the problem of finding *all* giant and dual-giant like solutions of a D3-brane in $AdS_5 \times S^5$ with scalars and electromagnetic fields on its world-volume turned on.

Result

- Our solutions can be elegantly described using the auxiliary space $\mathbb{C}^{1,2} \times \mathbb{C}^3$ with coordinates $\{\Phi^0, \Phi^1, \Phi^2; Z_1, Z_2, Z_3\}$ where the $AdS_5 \times S^5$ can be embedded as

$$|\Phi^0|^2 - |\Phi^1|^2 - |\Phi^2|^2 = l^2, \quad |Z_1|^2 + |Z_2|^2 + |Z_3|^2 = l^2.$$

Turning on gauge field does not change the embedding geometry of the D3-brane.

- So the D3-brane embeddings continue to be given by the Wobbling dual-giants and Mikhailove giants.

Wobbling dual-giants:

Defining $Y^i = \phi^i Z_1$, these are described as the intersection of

$$\begin{aligned} f(Y^0, Y^1, Y^2) &= 0 \quad \text{and} \\ Z_2 &= Z_3 = 0 \end{aligned} \tag{1}$$

with $AdS_5 \times S^5$.

Mikhailov giants:

These are the intersections of $AdS_5 \times S^5$ with

$$\begin{aligned} f(X^1, X^2, X^3) &= 0 \quad \text{and} \\ \phi^1 &= \phi^2 = 0, \end{aligned} \tag{2}$$

where $X^i = Z_i \phi^0$.

- Assuming that F is given by the pull-back of a bulk 2-form we find

For dual-giants

The susy world-volume gauge field F is given by the real part of

$$G = \sum_{ij=0,1,2} G_{ij}(Y) dY^i \wedge dY^j .$$

pulled back onto the world-volume.

For giants

The BPS world-volume gauge field strength is again given by the real part of

$$\tilde{G} = \sum_{ij=1,2,3} \tilde{G}_{ij}(X) dX^i \wedge dX^j , \quad (3)$$

pulled back onto the world-volume.

Computations

- To find supersymmetric configurations of a D-brane with bosonic fields it is sufficient to impose the κ -projection condition.
- We will solve the kappa-projection conditions a D3-brane embedded in $AdS_5 \times S^5$ with world-volume gauge field flux F .
- The metric on $AdS_5 \times S^5$ written in global coordinates is

$$\frac{ds^2}{l^2} = -\cosh^2 \rho d\phi_0^2 + d\rho^2 + \sinh^2 \rho (d\theta^2 + \cos^2 \theta d\phi_1^2 + \sin^2 \theta d\phi_2^2) \\ + d\alpha^2 + \sin^2 \alpha d\xi_1^2 + \cos^2 \alpha (d\beta^2 + \sin^2 \beta d\xi_2^2 + \cos^2 \beta d\xi_3^2)$$

where $\phi_0 = \frac{t}{l}$.

- This corresponds to parametrizing AdS_5 in $\mathbb{C}^{1,2}$ as

$$\begin{aligned}
 \phi^0 &= l \cosh \rho e^{i\phi_0}, \\
 \phi^1 &= l \sinh \rho \cos \theta e^{i\phi_1}, \\
 \phi^2 &= l \sinh \rho \sin \theta e^{i\phi_2}.
 \end{aligned}
 \tag{4}$$

- And S^5 in \mathbb{C}^3 is parametrized as

$$\begin{aligned}
 Z_1 &= l \sin \alpha e^{i\xi_1} \\
 Z_2 &= l \cos \alpha \sin \beta e^{i\xi_2} \\
 Z_3 &= l \cos \alpha \cos \beta e^{i\xi_3}
 \end{aligned}
 \tag{5}$$

- SUSY analysis needs an orthonormal frame for the $AdS_5 \times S^5$ metric ...

- We choose the following frame for the AdS_5 part of the metric

$$\begin{aligned}
 e^0 &= l[\cosh^2 \rho d\phi_0 - \sinh^2 \rho (\cos^2 \theta d\phi_1 + \sin^2 \theta d\phi_2)], \\
 e^1 &= l d\rho, \quad e^2 = l \sinh \rho d\theta, \\
 e^3 &= l \cosh \rho \sinh \rho (\cos^2 \theta d\phi_{01} + \sin^2 \theta d\phi_{02}) \\
 e^4 &= l \sinh \rho \cos \theta \sin \theta d\phi_{12}
 \end{aligned} \tag{6}$$

where $\phi_{ij} = \phi_i - \phi_j$.

- The base is the Kähler manifold $\widetilde{\mathbb{C}\mathbb{P}}^2$ along $\{r, \theta, \phi_{01}, \phi_{02}\}$, and the fibre along $\phi_0 + \phi_1 + \phi_2$.
- The Kähler form on $\widetilde{\mathbb{C}\mathbb{P}}^2$ is $\tilde{\omega} = e^{13} + e^{24}$.

- For the S^5 part we choose the frame

$$\begin{aligned}
 e^5 &= l d\alpha, & e^6 &= l \cos \alpha d\beta, \\
 e^7 &= l \cos \alpha \sin \alpha (\sin^2 \beta d\xi_{12} + \cos^2 \beta d\xi_{13}), \\
 e^8 &= l \cos \alpha \cos \beta \sin \beta d\xi_{23}, \\
 e^9 &= l (\sin^2 \alpha d\xi_1 + \cos^2 \alpha \sin^2 \beta d\xi_2 + \cos^2 \alpha \cos^2 \beta d\xi_3) \quad (7)
 \end{aligned}$$

where $\xi_{ij} = \xi_i - \xi_j$.

- This makes manifest the fact that S^5 is a Hopf fibration.
- The base is the Kähler manifold $\mathbb{C}P^2$ along $\{\alpha, \beta, \xi_{12}, \xi_{13}\}$ coordinates, and the fibre is along $\xi_1 + \xi_2 + \xi_3$.
- The Kähler form on $\mathbb{C}P^2$ is $\omega = e^{57} + e^{68}$.

- $AdS_5 \times S^5$ is a maximally susy solution of type IIB.
- The Killing spinor adapted to the above frame is

$$\epsilon = e^{-\frac{1}{2}(\Gamma_{79} - i\Gamma_5 \tilde{\gamma})\alpha} e^{-\frac{1}{2}(\Gamma_{89} - i\Gamma_6 \tilde{\gamma})\beta} e^{\frac{1}{2}\xi_1 \Gamma_{57}} e^{\frac{1}{2}\xi_2 \Gamma_{68}} e^{\frac{i}{2}\xi_3 \Gamma_9 \tilde{\gamma}} \\ \times e^{\frac{1}{2}\rho(\Gamma_{03} + i\Gamma_1 \gamma)} e^{\frac{1}{2}\theta(\Gamma_{12} + \Gamma_{34})} e^{\frac{i}{2}\phi_0 \Gamma_0 \gamma} e^{-\frac{1}{2}\phi_1 \Gamma_{13}} e^{-\frac{1}{2}\phi_2 \Gamma_{24}} \epsilon_0$$

- ϵ_0 is an arbitrary 32-component constant weyl spinor

$$\Gamma_0 \cdots \Gamma_9 \epsilon_0 = -\epsilon_0, \quad \gamma = \Gamma^{01234}, \quad \tilde{\gamma} = \Gamma^{56789}.$$

- We seek the full set of BPS equations for a single D3-brane preserving two susy out of the 32.
- We could choose the projections (not unique)

$$\begin{aligned} \Gamma_{57}\epsilon_0 = \Gamma_{68}\epsilon_0 &= i\epsilon_0, \\ \Gamma_{09}\epsilon_0 &= -\epsilon_0, \\ \Gamma_{13}\epsilon_0 = \Gamma_{24}\epsilon_0 &= -i\epsilon_0. \end{aligned}$$

- With these projections the killing spinor simplifies to

$$\epsilon = e^{\frac{i}{2}(\phi_0 + \phi_1 + \phi_2 + \xi_1 + \xi_2 + \xi_3)} \epsilon_0.$$

- We take the most general D3-brane ansatz ...
- ... all the space-time coordinates

$$(t, r, \theta, \phi_1, \phi_2, \alpha, \beta, \xi_1, \xi_2, \xi_3)$$

are functions of all the world-volume coordinates

$$(\sigma^0, \sigma^1, \sigma^2, \sigma^3).$$

- The kappa projection condition is

$$\epsilon^{ijkl} [\gamma_{ijkl} I + F_{ij} \gamma_{kl} I K + F_{ij} F_{kl} I] \epsilon = \sqrt{-\det(h + F)} \epsilon$$

[Cederwall, von Gussich, Nilsson, Westerberg; Bergshoeff, Townsend]

- Here, the world-volume gamma matrices are

$$\gamma_i = e_i^a \Gamma_a, \quad \text{where} \quad e_i^a = e_\mu^a \partial_i X^\mu,$$

with $i \in \{\tau, \sigma_1, \sigma_2, \sigma_3\}$ and $h_{ij} = \eta_{ab} e_i^a e_j^b$ is the induced metric.

- The operators K and I act as

$$K\epsilon = \epsilon^* \quad \text{and} \quad I\epsilon = -i\epsilon.$$

- Procedure:
 - ▶ Project the κ -equation on to some $\bar{\chi}_0$ and use the projection conditions to simplify it down to –
 - ▶ – a linear combination of non-vanishing and independent spinor bilinears $\bar{\chi}_0 \Gamma_{ab\dots} \epsilon_0$ and $\bar{\chi}_0 \Gamma_{ab\dots} \epsilon_0^*$.
 - ▶ Finally set their coefficients to zero.
- We have chosen to preserve the same susy as for the giant gravitons without electromagnetic flux.
- This means we have the condition

$$\gamma_{\sigma_0 \sigma_1 \sigma_2 \sigma_3} \epsilon = i \sqrt{-\det h} \epsilon$$

where $\gamma_{\sigma_0 \sigma_1 \sigma_2 \sigma_3} = e_0^a e_1^b e_2^c e_3^d \Gamma_{abcd}$.

- This leads to the constraint

$$\epsilon^{ijkl} F_{ij} \gamma_{kl} \epsilon_0^* = 0. \quad (8)$$

- This, in turn, gives rise to the condition

$$\sqrt{-\det h} + \text{Pf}[F] = \sqrt{-\det(h + F)}, \quad (9)$$

where $\text{Pf}[F] = \frac{1}{8} \epsilon^{ijkl} F_{ij} F_{kl}$ which we sometimes denote by “ $F \wedge F$ ”.

- To write the BPS equations in a compact form, define the complex one-forms

$$\mathbf{E}^1 = e^1 - ie^3, \quad \mathbf{E}^2 = e^2 - ie^4, \quad \mathbf{E}^5 = e^5 + ie^7, \quad \mathbf{E}^6 = e^6 + ie^8,$$

and the real 1-forms

$$\mathbf{E}^0 = e^0 + e^9 \quad \text{and} \quad \mathbf{E}^{\bar{0}} = e^0 - e^9$$

- It can be shown that the equation (8) linear in F gives

$$\begin{aligned}
 F \wedge \mathbf{E}^{AB} &= 0 \quad \text{for } A, B = \{1, 2, 5, 6\} \\
 F \wedge \mathbf{E}^0 \wedge \mathbf{E}^A &= 0 \quad \text{for } A = \{1, 2, 5, 6\} \\
 F \wedge (\mathbf{e}^{09} + i\Omega) &= 0,
 \end{aligned} \tag{10}$$

where $\Omega = \tilde{\omega} - \omega$, with

$$\begin{aligned}
 \tilde{\omega} &= \mathbf{e}^{13} + \mathbf{e}^{24} = -\frac{i}{2}(\mathbf{E}^{1\bar{1}} + \mathbf{E}^{2\bar{2}}) \quad \text{and} \\
 \omega &= \mathbf{e}^{57} + \mathbf{e}^{68} = \frac{i}{2}(\mathbf{E}^{5\bar{5}} + \mathbf{E}^{6\bar{6}}).
 \end{aligned}$$

- In these equations, by $F \wedge \mathbf{E}^{ab}$ we mean $\frac{1}{2}\epsilon^{ijkl} F_{ij} \mathbf{E}_k^a \mathbf{E}_l^b$ etc.
- Next we would like to solve equation (9).

- For this we note the following identity

$$\begin{aligned}
 -\det(h + F) &= -\det h - (\text{Pf}[F])^2 + (\epsilon^{09} \wedge F)^2 \\
 &\quad - \sum_{A=1,2,5,6} \left[|\epsilon^9 \wedge \mathbf{E}^A \wedge F|^2 - |\epsilon^0 \wedge \mathbf{E}^A \wedge F|^2 \right] \\
 &\quad - \sum_{A < B} |\mathbf{E}^{AB} \wedge F|^2 - (\Omega \wedge F)^2 + (\Omega \wedge \Omega) \text{Pf}[F].
 \end{aligned}$$

- Substituting the BPS conditions linear in the field strength into this expression, and noting that $\Omega \wedge \Omega = \frac{1}{2}(\tilde{\omega} - \omega) \wedge (\tilde{\omega} - \omega) = 0$ for time-like D3-branes (see [\[Ashok-NVS\]](#) for details) we obtain

$$\det(h + F) = \det h + (F \wedge F)^2. \quad (11)$$

Demanding the consistency of (9, 11) we get the last of the F -dependent BPS conditions

$$F \wedge F = 0. \quad (12)$$

This in turn implies $\det(h + F) = \det h$ for the BPS configurations we seek.

- Finally the BPS eqns that do not involve field strength F are:

$$\begin{aligned} \mathbf{E}^{ABCD} &= 0, & \mathbf{E}^{0ABC} &= 0 \\ (e^{09} + i\Omega) \wedge \mathbf{E}^{AB} &= 0 \text{ for } A, B, C, D = 1, 2, 5, 6 \\ \Omega \wedge \Omega &= 0. \end{aligned}$$

for time-like brane embeddings.

- In these equations we understand \mathbf{E}^{abcd} to be the function (0-form) $\epsilon^{ijkl} \mathbf{E}_i^a \mathbf{E}_j^b \mathbf{E}_k^c \mathbf{E}_l^d$ etc.
- Using all the BPS conditions, for a time-like D3-brane one obtains

$$\sqrt{-\det h} = e^{09} \wedge \Omega = i \mathbf{E}^{0\bar{0}} \wedge \sum_A \mathbf{E}^{A\bar{A}}. \quad (13)$$

- This is the “calibrating” form.

- For dual-giants, the BPS equations take the simplified form

$$\begin{aligned}\mathbf{E}^{0\bar{0}12} &= 0, \\ \mathbf{E}^0 \wedge \{\mathbf{E}^1, \mathbf{E}^2\} \wedge \tilde{\omega} &= 0, \\ \mathbf{E}^5 &= \mathbf{E}^6 = 0,\end{aligned}\tag{14}$$

while for giants, they take the form

$$\begin{aligned}\mathbf{E}^{0\bar{0}56} &= 0, \\ \mathbf{E}^0 \wedge \{\mathbf{E}^5, \mathbf{E}^6\} \wedge \omega &= 0, \\ \mathbf{E}^1 &= \mathbf{E}^2 = 0.\end{aligned}\tag{15}$$

- We now restrict our attention to dual-giant gravitons.

- Using the fact that the field-strength F is real, the F-dependent BPS conditions for the dual-giants take the form

$$\begin{aligned}
 F \wedge \mathbf{E}^{0\bar{0}} &= 0 & (16) \\
 F \wedge \mathbf{E}^0 \wedge \{\mathbf{E}^1, \mathbf{E}^2, \mathbf{E}^{\bar{1}}, \mathbf{E}^{\bar{2}}\} &= 0 \\
 F \wedge F &= 0 \\
 F \wedge (\mathbf{E}^{1\bar{1}} + \mathbf{E}^{2\bar{2}}) &= 0 \\
 F \wedge \{\mathbf{E}^{12}, \mathbf{E}^{\bar{1}\bar{2}}\} &= 0.
 \end{aligned}$$

- Next, we turn to solving these equations.

- At this point we make an assumption, that the field strength F on the world-volume can be written as a pull-back of a space-time 2-form onto the world-volume.
- This assumption allows us to solve the above algebraic conditions in a rather straightforward way.
- There are fifteen 2-forms that can be constructed out of the six bulk 1-forms of relevance $\{\mathbf{E}^0, \mathbf{E}^{\bar{0}}, \mathbf{E}^1, \mathbf{E}^2, \mathbf{E}^{\bar{1}}, \mathbf{E}^{\bar{2}}\}$ and the 2-form we seek is a real linear combination of all these two-forms.

- We start by assuming the most general ansatz for F :

$$F = \text{Re} \left[\chi_{0\bar{0}} \mathbf{E}^{0\bar{0}} + \sum_A (\chi_{0A} \mathbf{E}^{0A} + \chi_{\bar{0}A} \mathbf{E}^{\bar{0}A}) + \sum_{A \leq B} (\chi_{AB} \mathbf{E}^{AB} + \chi_{A\bar{B}} \mathbf{E}^{A\bar{B}}) \right]$$

- After using the BPS equations one will still be left with linear combinations of nine of the non-vanishing 4-forms on the left hand side of.
- We treat these nine 4-forms to be independent and set their coefficients to zero.
- This is justified because of our assumption that F can be written as the pull-back of a space-time 2-form irrespective of the details of the world-volume embedding equations.

- With this assumption it can be shown that the algebraic conditions can be solved if and only if

$$F = \text{Re}[\chi_{01}\mathbf{E}^{01} + \chi_{02}\mathbf{E}^{02} + \chi_{12}\mathbf{E}^{12}] \quad (17)$$

where $\chi_{01}, \chi_{02}, \chi_{12}$ are arbitrary complex functions of the bulk coordinates restricted to the world-volume.

- It now remains to solve for the parameters $\{\chi_{01}, \chi_{02}, \chi_{12}\}$ by imposing the Bianchi identity and the equation of motion for the gauge field.
- These are

$$dF = 0 \quad \text{and} \quad \partial_i X^{ij} = 0, \quad (18)$$

where

$$X^{ij} = \frac{1}{2} \sqrt{-\det(h + F)} [(h + F)^{-1} - (h - F)^{-1}]^{ij}. \quad (19)$$

- For any 4×4 symmetric matrix h whose components can be written as $h_{ij} = e_i^a e_j^b \eta_{ab}$ and for any antisymmetric 4×4 matrix F , we note the identity

$$\begin{aligned} & \det(h + F)[(h + F)^{-1} - (h - F)^{-1}]^{ij} \\ &= -\left(\frac{1}{4}\epsilon^{pqrs} F_{pq} F_{rs}\right) \epsilon^{ijmn} F_{mn} - \left(\frac{1}{2}\epsilon^{pqrs} F_{pq} e_r^a e_s^b\right) \eta_{ac} \eta_{bd} \epsilon^{ijmn} e_m^c e_n^d. \end{aligned}$$

- Given the definition of X^{ij} in and using the BPS equation $F \wedge F = 0$, we obtain

$$X^{ij} = \frac{1}{2\sqrt{-\det h}} \left(\frac{1}{2}\epsilon^{pqrs} F_{pq} e_r^a e_s^b\right) \eta_{ac} \eta_{bd} \epsilon^{ijmn} e_m^c e_n^d.$$

- We will need to simplify this further using the BPS equations.
- Before proceeding further we note that the equation of motion $\partial_i X^{ij} = 0$ can be written as $d\tilde{X} = 0$ for the 2-form defined as

$$\tilde{X} = \frac{1}{4} \epsilon_{ijmn} X^{mn} d\sigma^i \wedge d\sigma^j.$$

- We will therefore work with \tilde{X} and simplify it using our ansatz for the field strength F and the BPS equations.
- Substituting the ansatz we have for F and retaining only those terms which (potentially) survive after using the BPS equations one finds

$$\tilde{X} = -\frac{1}{\sqrt{-\det h}} \left[(F \wedge \mathbf{E}^{\bar{0}\bar{1}}) \mathbf{E}^{01} + (F \wedge \mathbf{E}^{\bar{0}\bar{1}}) \mathbf{E}^{0\bar{1}} \right. \\ \left. + (F \wedge \mathbf{E}^{\bar{0}\bar{2}}) \mathbf{E}^{02} + (F \wedge \mathbf{E}^{\bar{0}\bar{2}}) \mathbf{E}^{0\bar{2}} \right], \quad (20)$$

where we have re-expressed ϵ^{ab} in terms of \mathbf{E}^{ab} .

- It can be shown that when $F = \text{Re}[\chi_{01} \mathbf{E}^{01} + \chi_{02} \mathbf{E}^{02} + \chi_{12} \mathbf{E}^{12}]$

$$\tilde{X} = \frac{i}{\sqrt{-\det h}} (\mathbf{E}^{0\bar{0}1\bar{1}} + \mathbf{E}^{0\bar{0}2\bar{2}}) \text{Im}[\chi_{01} \mathbf{E}^{01} + \chi_{02} \mathbf{E}^{02} + \chi_{12} \mathbf{E}^{12}].$$

- For this we had to use the BPS equations and the identity

$$\mathbf{E}^{a[bcd} \mathbf{E}^{f]a} = 0,$$

where, as before, we understand \mathbf{E}^{abcd} to mean $\epsilon^{ijkl} \mathbf{E}_i^a \mathbf{E}_j^b \mathbf{E}_k^c \mathbf{E}_l^d$, and treat \mathbf{E}^{ab} as the rank-2 anti-symmetric tensor $\mathbf{E}_i^a \mathbf{E}_j^b - \mathbf{E}_j^a \mathbf{E}_i^b$.

- Finally restricting to the case of dual-giants, we have

$$\sqrt{-\det h} = i(\mathbf{E}^{0\bar{0}1\bar{1}} + \mathbf{E}^{0\bar{0}2\bar{2}}).$$

- Thus we finally obtain the result

$$\tilde{X} = \text{Im}[\chi_{01} \mathbf{E}^{01} + \chi_{02} \mathbf{E}^{02} + \chi_{12} \mathbf{E}^{12}]. \quad (21)$$

- This is a remarkable simplification, considering the original expression we started with, which was highly non-linear in the pull-back of the vielbeins and the field strength F .
- This can be attributed to the effectiveness of the BPS equations in simplifying the problem.
- Our final expressions for the real 2-forms F and \tilde{X} make it natural to define a complex 2-form

$$G = F + i\tilde{X} = \chi_{01}\mathbf{E}^{01} + \chi_{02}\mathbf{E}^{02} + \chi_{12}\mathbf{E}^{12} \quad (22)$$

in terms of which the Bianchi identity and the equations of motion can be combined into the single equation

$$dG = 0,$$

where dG refers to the exterior derivative of G on the world-volume.

- However, for differential forms, the pull-back operation and the exterior differentiation commute.
- So, we treat G as a spacetime 2-form and compute the exterior derivative in spacetime, and then require that the resulting 3-form vanishes, when pulled back onto the world-volume.
- Let us first recall some facts regarding the wobbling dual-giant solution. It is known that a wobbling dual-giant is described by a polynomial equation of the form

$$f(Y_0, Y_1, Y_2) = 0, \quad (23)$$

where $Y^i = \Phi^i Z_1$ with

$$\begin{aligned} \Phi^0 &= \cosh \rho e^{i\phi_0}, \quad \Phi^1 = \sinh \rho \cos \theta e^{i\phi_1}, \quad \Phi^2 = \sinh \rho \sin \theta e^{i\phi_2} \\ Z_1 &= e^{i\xi_1} \quad \text{and} \quad Z_2 = Z_3 = 0. \end{aligned} \quad (24)$$

- On such a 3 + 1 dimensional world-volume, we seek a closed complex 2-form of the kind (22).
- Since the equation of the D-brane is written purely in terms of the Y^i variables we rewrite (22) in terms of the differentials dY^i .
- Given the definition of the Y^i above, one can relate the differentials dY^i to the 1-forms E^A and $E^{\bar{A}}$. Using these one finds:

$$\begin{aligned}
 G &= \chi_{01} \mathbf{E}^{01} + \chi_{02} \mathbf{E}^{02} + \chi_{12} \mathbf{E}^{12} \\
 &:= G_{01} \frac{dY^0}{Y^0} \wedge \frac{dY^1}{Y^1} + G_{02} \frac{dY^0}{Y^0} \wedge \frac{dY^2}{Y^2} + G_{12} \frac{dY^1}{Y^1} \wedge \frac{dY^2}{Y^2} .(25)
 \end{aligned}$$

- Here the G_{ij} are given in terms of the χ_{ij} which can be inverted to express the χ_{ij} as linear combinations of G_{ij} since the matrix of coefficients is non-singular.

- The relations between the 2-form are

$$\begin{aligned}
 \frac{dY^0}{Y^0} \wedge \frac{dY^1}{Y^1} &= \frac{i}{\sinh \rho} \left[\frac{1}{\cosh \rho} E^{01} - \tan \theta E^{02} \right] - \frac{\tan \theta}{\cosh \rho} E^{12}, \\
 \frac{dY^0}{Y^0} \wedge \frac{dY^2}{Y^2} &= \frac{i}{\sinh \rho} \left[\frac{1}{\cosh \rho} E^{01} + \cot \theta E^{02} \right] + \frac{\cot \theta}{\cosh \rho} E^{12}, \\
 \frac{dY^1}{Y^1} \wedge \frac{dY^2}{Y^2} &= \frac{1}{\sinh \rho \cos \theta \sin \theta} \left[iE^{02} + \coth \rho E^{12} \right]. \quad (26)
 \end{aligned}$$

- These combinations of $\{E^{01}, E^{02}, E^{12}\}$ have the important property that their exterior derivatives vanish - a useful fact.

- Let us now turn to solving the equation $dG = 0$. The left hand side of this equation reads

$$\begin{aligned}
 dG = & dG_{01} \wedge \left(\frac{i}{\sinh \rho} \left[\frac{1}{\cosh \rho} \mathbf{E}^{01} - \tan \theta \mathbf{E}^{02} \right] - \frac{\tan \theta}{\cosh \rho} \mathbf{E}^{12} \right) \\
 & + dG_{02} \wedge \left(\frac{i}{\sinh \rho} \left[\frac{1}{\cosh \rho} \mathbf{E}^{01} + \cot \theta \mathbf{E}^{02} \right] + \frac{\cot \theta}{\cosh \rho} \mathbf{E}^{12} \right) \\
 & + dG_{12} \wedge \left(\frac{1}{\sinh \rho \cos \theta \sin \theta} \left[i \mathbf{E}^{02} + \coth \rho \mathbf{E}^{12} \right] \right), \quad (27)
 \end{aligned}$$

with

$$\begin{aligned}
 dG_{ij} = & (K_0 G_{ij}) \mathbf{E}^0 + (K_1 G_{ij}) \mathbf{E}^1 + (K_2 G_{ij}) \mathbf{E}^2 + (K_{\bar{0}} G_{ij}) \mathbf{E}^{\bar{0}} \\
 & + (K_{\bar{1}} G_{ij}) \mathbf{E}^{\bar{1}} + (K_{\bar{2}} G_{ij}) \mathbf{E}^{\bar{2}},
 \end{aligned}$$

where K_A is the vector field dual to the 1-form E^A .

- Now, (27) is an equation for a 3-form on the world-volume of the dual-giant and one should set the coefficients of the linearly independent 3-forms to zero. As before, we will do this pretending that this is a bulk 3-form equation given the form of our ansatz. Such a solution would lead to a spacetime 2-form G , independent of the particular polynomial that defines the dual-giant. We implement this procedure below.
- Given that the equation describing the dual-giant is holomorphic in the variables Y^i , it follows that $\mathbf{E}^{012} = 0$. This is true irrespective of the precise form of the defining polynomial $f(Y^i) = 0$.
- Also, from (27), it follows that two of the three indices in the 3-form have to be unbarred. Given these constraints, there are precisely nine independent 3-forms that appear on the right hand side of that equation if we substitute into (27).

- After some algebra we find that imposing $dG = 0$ is equivalent to imposing the nine constraints

$$K_{\bar{0}}G_{ij} = K_{\bar{1}}G_{ij} = K_{\bar{2}}G_{ij} = 0 \quad \text{for } i, j \in \{0, 1, 2\}. \quad (28)$$

- These make G_{ij} to be functions of Y^0, Y^1, Y^2 and not their conjugates.
- We can summarize our results so far as follows:

- Any 1/8-BPS dual-giant in $AdS_5 \times S^5$ with non-trivial world-volume electromagnetic fields is specified by
 - ▶ a holomorphic function

$$f(Y^0, Y^1, Y^2) \quad \text{and}$$

- ▶ a holomorphic 2-form

$$G = \sum_{i,j=0,1,2} G_{ij} \frac{dY^i}{Y^i} \wedge \frac{dY^j}{Y^j},$$

with $G_{ij} = G_{ij}(Y^0, Y^1, Y^2)$.

- The world-volume is obtained by taking the intersection of the zero-set of $f(Y^0, Y^1, Y^2)$ with $AdS_5 \times S^5$ and the field strength of the world-volume gauge field is given by the real part of the 2-form G pulled back onto the world-volume.
- This completes our general analysis of the BPS equations for dual-giants in the $AdS_5 \times S^5$ background of type IIB supergravity.
- Similar analysis gives rise to the corresponding result for giants.

Conclusion

- We have (with an assumption) generalized all the Mikhailov giants and Wobbling dual-giants to include the world-volume gauge field.
- We have a world-volume reparametrization invariant description.
- What about fermions?
- What about charges and quantization?